



Discrete particle simulations with MercuryDPM

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15 May 2025



A bit about the instructors



Co-founded MercuryDPM in 2009



Main Developer/Manager



A bit about the instructors



Co-founded MercuryDPM in 2009



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Particle simulations with MercuryDPM



Particle simulations with MercuryDPM



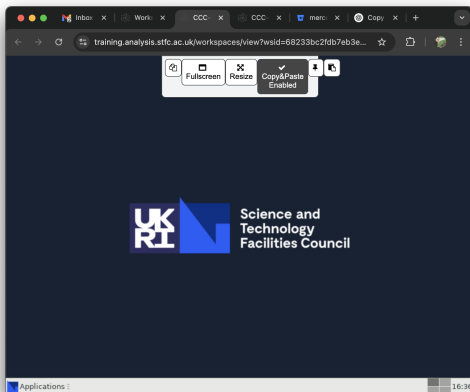
Open-source software for simulating granular materials

- Developed since 2009
- Modern C++ code base, easily extendable
- Funded by many research projects (>€10 mil.)
- Worldwide user/developer base
- Available at mercurydpm.org



But, how?

- 1 Go to <https://training.analysis.stfc.ac.uk/workspaces/>
- 2 Open workspace CCC-ParaSolS Training (CPU Only)
- 3 Hover on top of the window to get the menu below, and enable Copy&Paste





Exercise 0: Let's install - In 5 easy steps

- 1 Open a terminal
- 2 Create a folder
`mkdir MercuryDPM`
`cd MercuryDPM`
- 3 Download the code:
`git clone https://bitbucket.org/mercurydpm/mercurydpm.git MercurySource`
- 4 Create the build system:
`mkdir MercuryBuild`
`cd MercuryBuild`
`cmake ../MercurySource`
- 5 Build and test the code: `make fullTest`

Full instructions here: <https://www.mercurydpm.org/downloads>



Meanwhile while installing ...



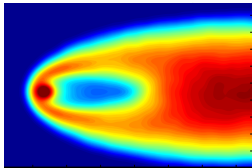
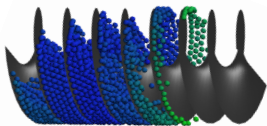
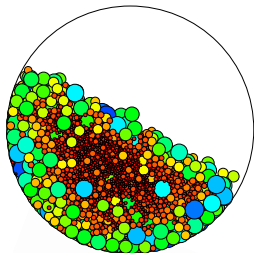
Directory Structure

MercurySource contains many subdirectories:

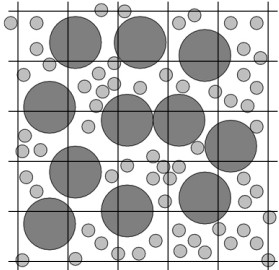
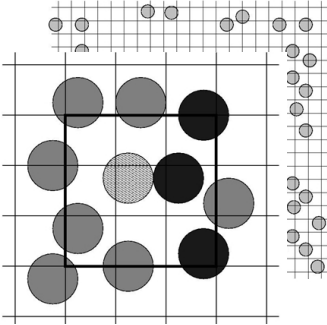
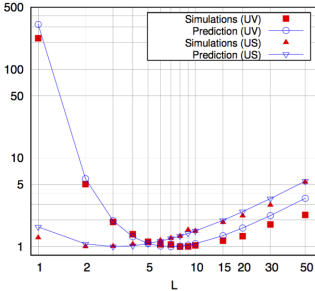
- **Kernel** contains all core functionality of MercuryDPM:
Classes to store and manipulate particles, wall, boundaries, contact laws (species), and the DPM algorithm (**Mercury3D**).
- **Drivers** contains codes to simulate many different processes, tests, and demos.
These codes use the classes defined in the Kernel.
- **Documentation** Text files and figures required to build the documentation
- **Configuration** Config files, e.g. for doxygen
- **Tools** Scripts for postprocessing and visualisation
- **Scripts** Matlab/Python scripts for postprocessing and testing
- **XBalls** A simple visualisation software

Particle simulations with MercuryDPM

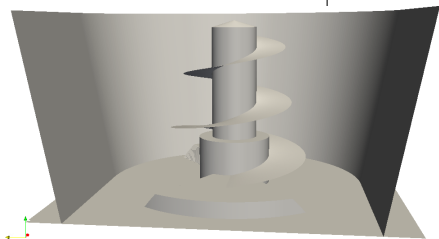
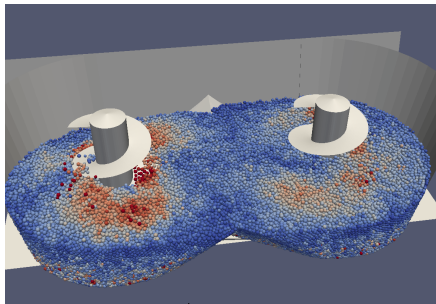
- **Fast**
Contact detection algorithm
allows polydisperse simulations
- **Flexible**
Complex walls and
boundary conditions
- **Accurate**
Coarse-graining technique
to extract continuum fields
- **Open-source**
Available at mercurydpm.org



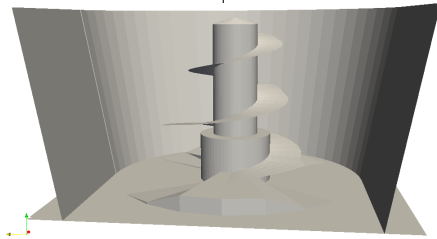
Fast computations (optimal contact detection)

Traditional Linked Cell	MercuryDPM Hierarchical Grid	Speed-Up for high polydispersity
 #checks: 1000	 #checks: 89	 up to 200x quicker!

Curved walls: flexible, efficient and accurate

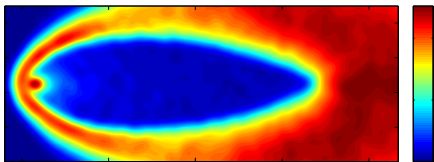
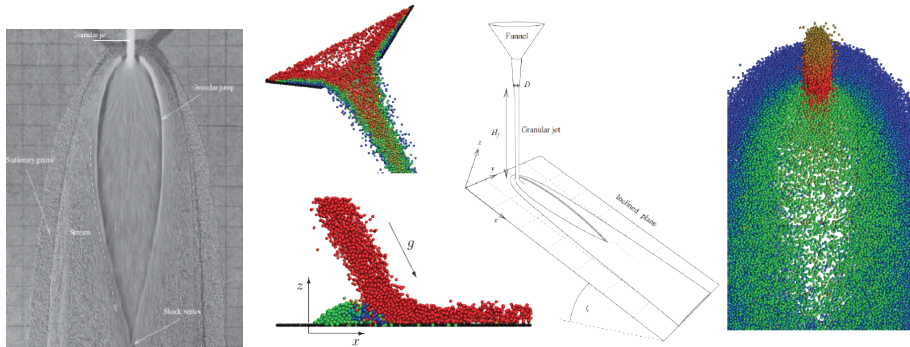


Curved walls: 5 surfaces

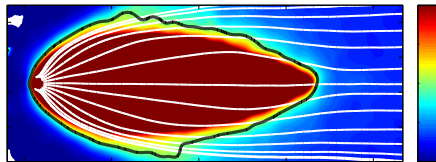


STL: 1456 surfaces

Accurate analysis via MercuryCG



flow height

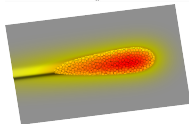
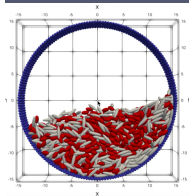
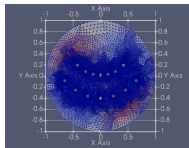


shocks & streamlines



Other key features

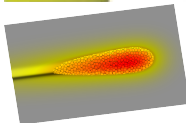
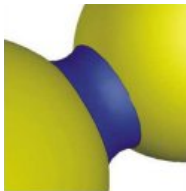
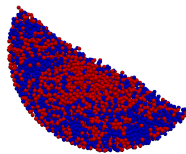
- Integration with fluid/solid solvers
- Fully parallelised
- Non-spherical particles
- Model complex physical interactions
- Model many complex processes





Other key features

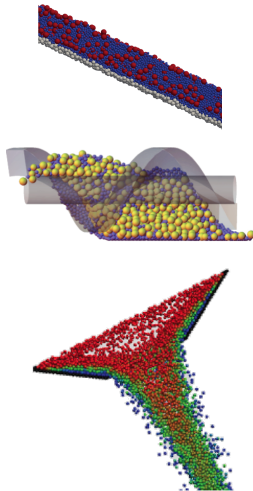
- Integration with fluid/solid solvers
- Fully parallelised
- Non-spherical particles
- Model complex physical interactions
 - Cohesion/adhesion
 - Liquid bridges
 - Charged particles
 - Sintering/melting
 - Surface wear
 - Breakage
- Model many complex processes





Other key features

- Integration with fluid/solid solvers
- Fully parallelised
- Non-spherical particles
- Model complex physical interactions
- Model many complex processes
 - Tabletting/agglomeration
 - Selective laser sintering
 - Cohesive mixing, granulation
communion and deagglomeration
 - Conveying/transportation
 - Spray-drying





External Packages

- Git is used to download the code and get updates.
- CMake for compiling, CTest for self-testing.
- Python for self-testing and pre/postprocessing.
- Doxygen for documentation.
- X11 for xballs visualisation.
- Paraview for 3D visualisation.
- MPI for running parallelised simulations.



Git

MercuryDPM is stored in a git repository.

- Download the code:
`git clone https://bitbucket.org/mercurydpm/mercurydpm.git`
- Update to the latest version: `git pull`
- Check for local modifications: `git status`
- Get other branches of the code: `git checkout 1.0.Alpha`

Why Git/Bitbucket?

Git gives version history, allows parallel development.

Bitbucket allows us to share the code open-source.



CMake

- All subfolders containing source files also contain a CMakeLists.txt file to define how the files should be compiled and linked together.
- Running **cmake-gui**, **ccmake** or **cmake** will create a build directory, that contains Makefiles to build the executables.
- Important: Rerun **cmake** each time you create a new source file.

Why CMake?

Using CMake, we tell the compiler how to create executables from the source files.



CMake

- CMake creates a build directory (**MercuryBuild**), which mirrors the structure of the source directory (**MercurySource**).
 - Source directory
 - Drivers
 - Kernel
 - CMakeLists.txt
 - ...
 - Build directory
 - Drivers
 - Kernel
 - CMakeCache.txt
 - ...
- Use **cmake .** to update the build instructions, e.g. when you add new files.
- Use **make name** to build the executable belonging to the driver **name.cpp**.
- Special make commands : **make fullTest** runs the test suite, **make clean** removes all compiled files, **make doc** builds the documentation.



CTest

- All drivers labelled `*UnitTest.cpp`, `*SelfTest.cpp`, `*MPI*Test.cpp` are part of the test suite. Every time the code is updated, we test whether these codes still produce the desired output.
- `make fullTest` runs the full test suite (clears, compiles, and runs all tests), `make test` runs all tests in a directory.
- If a test fails, it produces an errorlog (e.g. `STLReaderUnitTest.errorlog`).

Why CTest?

CTest allows us to check whether the current development has affected the existing code. This ensures that modifications don't break existing features.



Doxygen Documentation

- Uses doxygen to document all functions and classes in MercuryDPM
- Includes tutorials and demos, output files, visualisation, how to develop a new driver, how to use MPI, etc.
- Can be built locally with **make doc** (if turned ON in cmake)
- Web version: <http://docs.mercurydpm.org/Trunk>

MercuryDPM Trunk

Main Page Related Pages Namespaces Classes Files Search

MercuryDPM

- MercuryDPM Reference Manual
 - Installing MercuryDPM
 - Getting started
 - Fun with MercuryDPM (Tutorials)
 - Write Your Own Code
 - Coarse graining: From discrete parti
 - Visualising Your Results
 - Guidelines for Users and Developer
 - Contact Models in MercuryDPM
 - For Developers
 - Todo List
 - Deprecated List
 - Bug List
 - Namespaces
 - Classes
 - Files

Getting started

- Introduction to the Code
- Basic classes in MercuryDPM
- Directory Structure

MercuryDPM Reference Manual

Generated on Wed Jul 17 2019 04:00:23 for MercuryDPM by doxygen 1.8.7



Editors

IDEs

- Not just an editor, but a full-scale “development environment”
- Many features that usually operates on a “whole project”
 - Loads the project
 - Allows navigation between files
 - Provides autocompletion based on the whole project
 - Integrates with a version management system (like git)
 - A complete testing environment
- CLion, Eclipse, Visual Studio

Lightweight Editors

- Not as powerful as IDEs, but they’re fast, elegant and simple
- They are mainly used to open and edit a file instantly
- However, in practice, lightweight editors may have a lot of plugins similar to IDEs, so there’s no strict border between a lightweight editor and an IDE
- Sublime Text, NotePad++, Vim and Emacs



Let's crack on with real exercises!!!



Exercise 1: Build the documentation

- Go to MercuryBuild.
- Use cmake (or ccmake) to turn on the option of building the documentation:
`cmake ../MercurySource -DMercuryDPM_BUILD_DOCUMENTATION=ON`
- Type `make doc` to create a local documentation
- Open the documentation in a web browser
`firefox Documentation/html/index.html &`



Exercises: Tutorials 1-3

- Goto **MercuryBuild** and open the documentation in a web browser
`firefox Documentation/html/index.html` &
- Click on "Fun with MercuryDPM (Tutorials)".
- Then, click on "Beginner tutorials".
- Read, execute and visualise the first three tutorials.



Deep dive into Tutorial 3



Output files

- **Tutorial3.data**
contains particle positions, velocities, radii for multiple time steps
- **Tutorial3.fstat**
contains contact points, forces, overlaps for multiple time steps
- **Tutorial3.ene**
contains kinetic and potential energy, center of mass for multiple time steps
- **Tutorial3.restart**
contains all parameters of the last saved time step.



fileName.data

- First line: N, Time, XMin, YMin, ZMin, XMax, YMax, ZMax
- N_p lines: x, y, z, vx, vy, vz, rad, q1, q2, q3, ox, oy, oz, species
- ... then repeats

```
1 0 0 0 0.01 0.1
0.005 0.095 0 0 0.005 -0 -0 0
1 0.0005 0 0 0.01 0.1
0.005 0.094998775 0 -0.0049 0.005 -0 -0 0
...
```

Used for visualisation and analysis.



fileName.fstat

- 3 lines: # time, info; # info; # info
- N_c lines: $t, i, j, x, y, z, \delta, \delta^t, f^{c,n}, f^{c,t}, n_x, n_y, n_z, t_x, t_y, t_z$
- ... then repeats

```
# 0 1
# 0 0 0 169.9 169.9 169.9
# 1 40 0 0 0 0
0 5 6 43.93 6.894 134.1 1.360 0 0 0 0.9995 -0.009 -0.0274 0 0 0
0 6 5 43.93 6.894 134.1 1.360 0 0 0 -0.9995 0.009 0.0274 0 0 0
...
```

Used for analysis.



fileName.ene

- Stores energy statistics.

t	ene_{gra}	ene_{kin}	ene_{rot}	ene_{ela}	X_{COM}	Y_{COM}	Z_{COM}
0	0.0009..	0	0	0	0.005	0.095	0
0.0005	0.0009..	$1.2e - 08$	0	0	0.005	0.094	0
...							

Used for quickly checking whether the program runs as expected.

fileName.restart

- Stores the full state of the simulation.
- Used for restarting, checking parameters.

```
MercuryDPM 0.14 name Tutorial3 revision 4946 repository https://mercurydpm.  
dataFile      fileType MULTIPLE_FILES saveCount 10 counter 2000  
fStatFile     fileType NO_FILE saveCount 10 counter 0  
eneFile       fileType ONE_FILE saveCount 10 counter 2000  
restartFile   fileType ONE_FILE saveCount 10 counter 2000  
xMin 0 xMax 1 yMin 0 yMax 1 zMin 0 zMax 1  
timeStep 0.0001 time 2 ntimeSteps 20000 timeMax 2  
systemDimensions 3 particleDimensions 3 gravity 0 0 0 writeVTK 0  
Species 1  
LinearViscoelasticSpecies id 0 density 2500 stiffness 258.5 dissipation 0  
Walls 0  
Boundaries 0  
Particles 1  
BaseParticle id 0 indSpecies 0 position 1 0.3 0.3 orientation 1 0 0 0 veloc  
Interactions 0
```

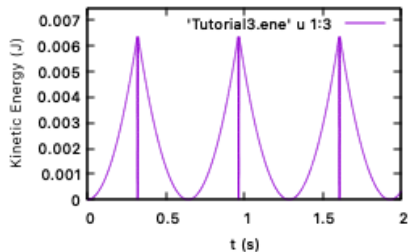
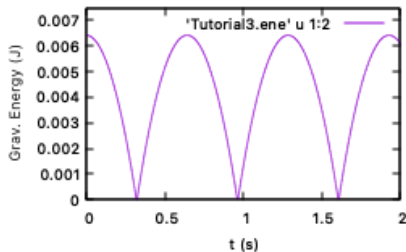
T3 exercise solution

② Open **gnuplot**

③ Plot gravitational and kinetic energy:

`plot 'Tutorial3.ene' using 1:2 with lines`

`plot 'Tutorial3.ene' using 1:3 with lines`¹



We observe a small collision time (5ms), and no loss of velocity ($r=1.0$).

¹You need to reduce savecount to 10 to get a detailed kinetic energy plot



Typical driver code



A typical driver code

```
#include "Mercury3D.h"

class Demo : public Mercury3D {
    void setupInitialConditions() override {
        //define geometric setup here
    }
};

int main() {
    Demo problem;
    //define process parameters here
    problem.solve();
}
```

- **Mercury3D** contains the DPM algorithm.
- The derived class **Demo** contains the initial conditions.
- The main function instantiates the class and calls the solve routine.



A typical driver code: the **main** function

Define parameters in the main function using set-functions:

```
...  
int main() {  
    Demo problem;  
    problem.setName("Demo");  
    problem.setMin(Vec3D(-1.0, -1.0, -1.0));  
    problem.setMax(Vec3D(1.0, 1.0, 1.0));  
    problem.setTimeStep(1e-4);  
    problem.setTimeMax(2.0);  
    problem.setGravity(Vec3D(0, 0, -9.81));  
    problem.setSaveCount(200);  
    problem.solve();  
}
```



A typical driver code: `setupInitialConditions`

- Define the initial state in `setupInitialConditions`.
- You can declare four main types of objects:
 - Particles
 - Walls
 - Species (material and contact properties)
 - Boundaries (boundary conditions like periodic walls, insertion and deletion regions)
- Include the appropriate class, declare a new object, set its parameters and add it to the handler.

Let's look at a few examples ...



A typical driver code

First, define a species (material and contact properties):

```
#include "Species/LinearViscoelasticSpecies.h"
...
void setupInitialConditions()
{
    ...
    LinearViscoelasticSpecies species;
    species.setDensity(2000);
    species.setStiffness(1e3);
    species.setDissipation(0.1);
    speciesHandler.copyAndAddObject(species);
    ...
}
...
```



A typical driver code

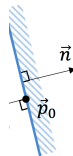
Then, define particles:

```
#include "Particles/SphericalParticle.h"
...
void setupInitialConditions()
{
    ...
    SphericalParticle p;
    p.setSpecies(speciesHandler.getObject(0));
    p.setRadius(1e-3);
    p.setPosition(Vec3D(0.1,0.1,0.0));
    particleHandler.copyAndAddObject(p);
    ...
}
...
```



A typical driver code

Now add walls by defining a wall's normal and position.



```
#include "Walls/InfiniteWall.h"
...
void setupInitialConditions()
{
    ...
    InfiniteWall w0;
    w0.setSpecies(speciesHandler.getObject(0));
    w0.set(Vec3D(0.0, 0.0, -1.0), Vec3D(0, 0, getZMin()));
    wallHandler.copyAndAddObject(w0);
    ...
}
...
```



A typical driver code

Note: for MPI (parallel) simulations, you have to define the species in the main:

```
#include "Species/LinearViscoelasticSpecies.h"
...
int main() {
    ...
    LinearViscoelasticSpecies species;
    species.setDensity(2000);
    species.setStiffness(1e3);
    species.setDissipation(0.1);
    problem.speciesHandler.copyAndAddObject(species);
    ...
}
```

For an instructive full example, see `Tutorial3_BouncingBallElastic.cpp`.



Tutorials T4 -T9 exercises



Come join the team :)

<https://www.mercurydpm.org/about/team>



Behind the scenes:

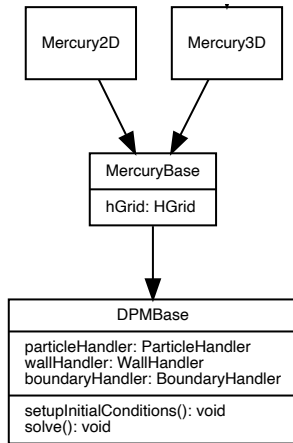
The MercuryDPM kernel



MercuryDPM's class structure

The main class of MercuryDPM is **Mercury3D**.

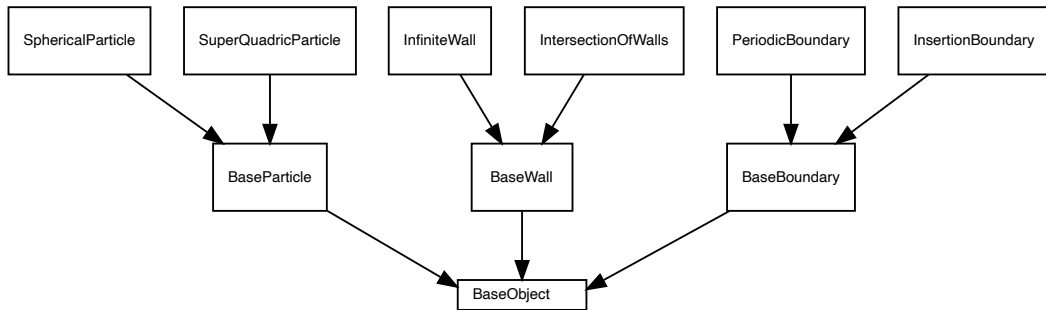
We use a cascade of classes to successively add complexity to the DPM algorithm:





MercuryDPM's class structure

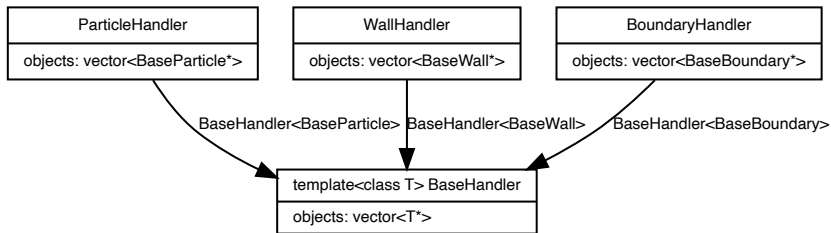
All other objects are also sucessively derived from simpler classes:





MercuryDPM's class structure

All objects are stored in handlers:



With that information, the driver codes are easier to understand.



The function `solve()`

The solve routine in `DPMBase.cc` is the work horse of the code. A simplified version:

```
void DPMBase::solve()
{
    if (isRestarted()) {
        actionsOnRestart();          // user-defined actions
    } else {
        setupInitialConditions();    // user-defined actions
    }

    while (getTime() < getTimeMax()) {
        computeOneTimeStep();
    }
}
```



The function `computeOneTimeStep()`

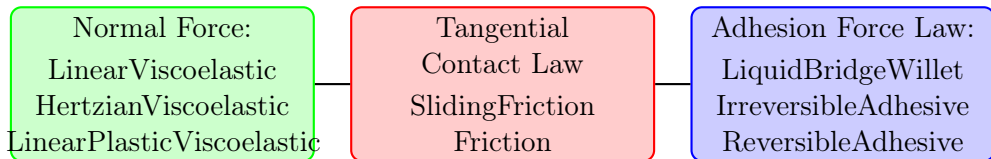
`computeOneTimeStep` evolves the system by one timestep. It also provides action-functions that can be specified by the user. A simplified version:

```
void DPMBase::computeOneTimeStep()
{
    integrateBeforeForceComputation(); // velocity-verlet part 1
    actionsBeforeTimeStep();           // user-defined actions
    computeAllForces();                 // force computation
    actionsAfterTimeStep();             // user-defined actions
    integrateAfterForceComputation();   // velocity-verlet part 2
    time_ += timeStep_;
}
```



Material and contact properties

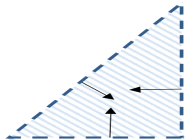
- **Species** contain material and contact properties.
- Material properties: density, etc
- Contact properties: stiffness, dissipation coefficient, etc
- Species can combine up to three different contact laws:



- The species name is a concatenation of the three force laws, e.g.
LinearViscoelasticFrictionLiquidBridgeWilletSpecies
- For more about the various force laws implemented in MercuryDPM, visit
<https://doi.org/10.1016/j.cpc.2019.107129>



Polygonal wall



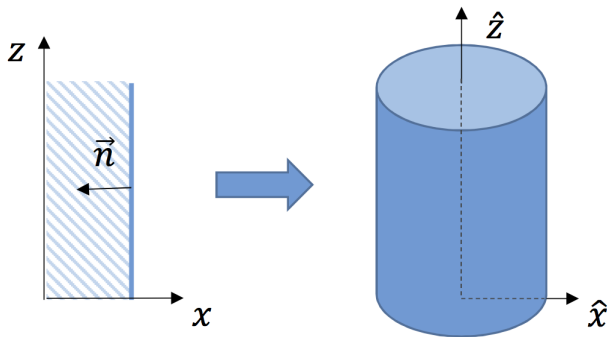
- Constructed by intersection of infinite walls.

```
#include "Walls/IntersectionOfWalls.h"
...
void setupInitialConditions()
{
    IntersectionOfWalls w;
    w.addObject(Vec3D(1,-1,0),Vec3D(0,0,0));
    w.addObject(Vec3D(1, 0,0),Vec3D(0,0,0));
    w.addObject(Vec3D(0,-1,0),Vec3D(1,0,0));
    wallHandler.copyAndAddObject(w);
}
...
```



Axisymmetric walls

Obtained by rotating a 2D-polygonal wall around an axis:



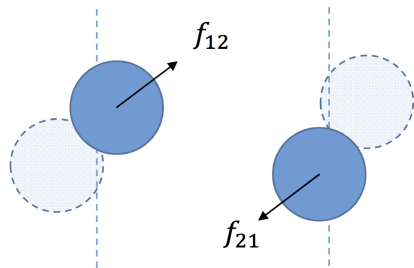


Axisymmetric walls

```
#include "Walls/IntersectionOfWalls.h"
...
void setupInitialConditions()
{
    AxisymmetricIntersectionOfWalls w;
    w.setPosition(0.,0.,0.);
    w.setOrientation(Vec3D(0.0,0.0,1.0));
    w.addObject(Vec3D(1,0,0), Vec3D(2.0,0.0,0.0));
    wallHandler.copyAndAddObject(w);
}
...
```

See HourGlass3DSelfTest in MercurySimpleDemos.

Periodic boundaries



- Found in Kernel/Boundaries/.
- See documentation.
- Similar: AngledPeriodicBoundary
- Exercise: Modify **FreeCoolingDemo.cpp** to use periodic boundaries in x - and y -direction.

```
#include "Boundaries/PeriodicBoundary.h"
...
void setupInitialConditions()
{
    PeriodicBoundary b;
    b.set(Vec3D(1,0,0), getXMin(), getXMax());
    boundaryHandler.copyAndAddObject(b);
}
...
```



LeesEdwardsBoundary

Simulates a shear box, see LeesEdwardsSelfTest.cpp.

```
double velocity = 1.0;
auto posShift = [velocity] (double time) {return time*velocity;};
auto velShift = [velocity] (double time) {return velocity;};
LeesEdwardsBoundary b;
b.set(posShift, velShift, getXMin(), getXMax(), getYMin(), getYMax());
boundaryHandler.copyAndAddObject(b);
```



InsertionBoundaries

- Inserts particles in a volume, e.g. cube.
- By default, inserts until no more particles can be inserted (maxFailed criterium).
- Use `setPSD(..)` to set a particle size distribution.
- `setInitialVolume(..)`, `setVolumeFlowRate(..)` to insert a defined amount.
- Change the defaultParticle properties to determine what gets inserted.

```
SphericalParticle defaultParticle;  
p.setSpecies(speciesHandler.getObject(0));
```

```
CubeInsertionBoundary b;  
b.set(&defaultParticle, maxFailed, posMin, posMax, velMin, velMax, radMin,  
    ↪ radMax);  
boundaryHandler.copyAndAddObject(b);
```



DeletionBoundaries

- Deletes particles in a volume.

```
DeletionBoundary b;  
b.set(Vec3D(0, 0, -1), -minHeight);  
boundaryHandler.copyAndAddObject(b);
```





Fixed particles, prescribed position

- Particles and walls can have a prescribed position/velocity/angular velocity
- To create rough surfaces

```
SphericalParticle p;  
p.fixParticle();  
...
```

```
auto position = [] (double t)  
{ return Vec3D(0, 0, std::sin(t)); };  
InfiniteWall w;  
w.setPrescribedPosition(position);  
...
```





Parameter studies

- `autonumber()` saves files with an appended number.
- The number can be used to run parameter studies.
- Set up a 4-by-4 parameter study:

```
int main() {  
    ...  
    problem.autoNumber();  
    std::vector<int> study_num=problem.get2DParametersFromRunNumber(4,4);  
    double parameter1 = 2.0 + (study_num[1] - 1.) * 0.5;  
    double parameter2 = (0.2 + 0.1 * study_num[2]);  
}
```



Further reading

- Lots of resources here: <https://www.mercurydpm.org/documentation>
- Including the text book: <https://oercommons.org/courseware/lesson/113895>
- Or the source for the must up-to version are here:
<https://bitbucket.org/mercurydpm/scientificcomputingcourse/src/master/>
- This includes the version for this course



Further reading

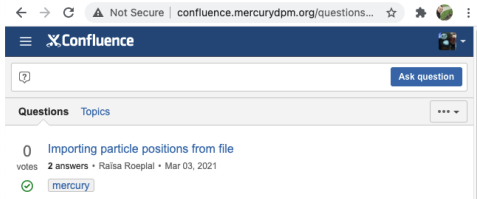
- This journal article gives a good overview of the implementation and features:
<https://doi.org/10.1016/j.cpc.2019.107129>
- If you have questions, please visit our Q&A:
<https://mercurydpm.atlassian.net/wiki/spaces/CQ/overview/>



CPC 50th anniversary article

Fast, flexible particle simulations – An introduction to MercuryDPM^{a,b,*,☆}

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<https://www.mercurydpm.org/about/team>